

# FUNDAMENTALS OF OPTICS Fourth Edition

---

**Francis A. Jenkins**  
*Late Professor of Physics*  
*University of California, Berkeley*

**Harvey E. White**  
*Professor of Physics, Emeritus*  
*Director of the Lawrence Hall of Science, Emeritus*  
*University of California, Berkeley*

 **McGraw-Hill Primis  
Custom Publishing**

New York St Louis San Francisco Auckland Bogotá  
Caracas Lisbon London Madrid Mexico Milan Montreal  
New Delhi Paris San Juan Singapore Sydney Tokyo Toronto

teristic of any interference experiment with light that the sources must have this point-to-point phase relation, and sources that have this relation are called *coherent sources*.

While special arrangements are necessary for producing coherent sources of light, the same is not true of *microwaves*, which are radio waves of a few centimeters wavelength. These are produced by an oscillator which emits a continuous wave, the phase of which remains constant over a time long compared with the duration of an observation. Two independent microwave sources of the same frequency are therefore coherent and can be used to demonstrate interference. Because of the convenient magnitude of their wavelength, microwaves are used to illustrate many common optical interference and diffraction effects.\*

If in Young's experiment the source slit  $S$  (Fig. 13C) is made too wide or the angle between the rays which leave it too large, the double slit no longer represents two coherent sources and the interference fringes disappear. This subject will be discussed in more detail in Chap. 16.

### 13.8 DIVISION OF AMPLITUDE. MICHELSON† INTERFEROMETER

Interference apparatus may be conveniently divided into two main classes, those based on *division of wave front* and those based on *division of amplitude*. The previous examples all belong to the former class, in which the wave front is divided laterally into segments by mirrors or diaphragms. It is also possible to divide a wave by partial reflection, the two resulting wave fronts maintaining the original width but having reduced amplitudes. The Michelson interferometer is an important example of this second class. Here the two beams obtained by amplitude division are sent in quite different directions against plane mirrors, whence they are brought together again to form interference fringes. The arrangement is shown schematically in Fig. 13N. The main optical parts consist of two highly polished plane mirrors  $M_1$  and  $M_2$  and two plane-parallel plates of glass  $G_1$  and  $G_2$ . Sometimes the rear side of the plate  $G_1$  is lightly silvered (shown by the heavy line in the figure) so that the light coming from the source  $S$  is divided into (1) a reflected and (2) a transmitted beam of equal intensity. The light reflected normally from mirror  $M_1$  passes through  $G_1$  a third time and reaches the eye as shown. The light reflected from the mirror  $M_2$  passes back through  $G_2$  for the second time, is reflected from the surface of  $G_1$  and into the

\* The technique of such experiments is discussed by G. F. Hull, Jr., *Am. J. Phys.*, 17:599 (1949).

† A. A. Michelson (1852–1931). American physicist of genius. He early became interested in the velocity of light and began experiments while an instructor in physics and chemistry at the Naval Academy, from which he graduated in 1873. It is related that the superintendent of the Academy asked young Michelson why he wasted his time on such useless experiments. Years later Michelson was awarded the Nobel prize (1907) for his work on light. Much of his work on the speed of light (Sec. 19.3) was done during 10 years spent at the Case Institute of Technology. During the latter part of his life he was professor of physics at the University of Chicago, where many of his famous experiments on the interference of light were done.

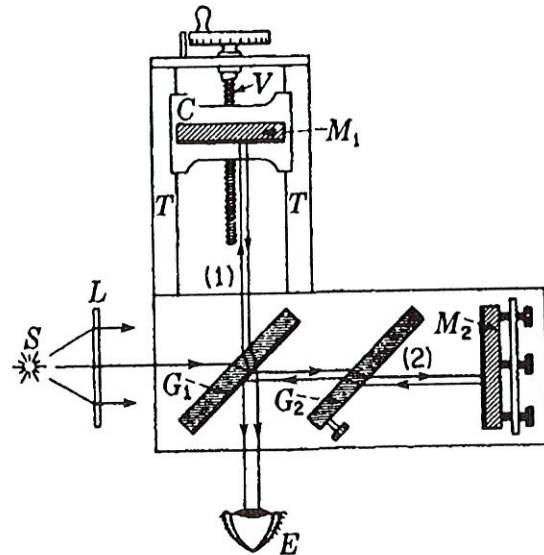


FIGURE 13N  
Diagram of the Michelson interferometer.

eye. The purpose of the plate  $G_2$ , called the *compensating plate*, is to render the path *in glass* of the two rays equal. This is not essential for producing fringes in monochromatic light, but it is indispensable when white light is used (Sec. 13.11). The mirror  $M_1$  is mounted on a carriage  $C$  and can be moved along the well-machined ways or tracks  $T$ . This slow and accurately controlled motion is accomplished by means of the screw  $V$ , which is calibrated to show the exact distance the mirror has been moved. To obtain fringes, the mirrors  $M_1$  and  $M_2$  are made exactly perpendicular to each other by means of screws shown on mirror  $M_2$ .

Even when the above adjustments have been made, fringes will not be seen unless two important requirements are fulfilled. First, the light must originate from an *extended* source. A point source or a slit source, as used in the methods previously described, will not produce the desired system of fringes in this case. The reason for this will appear when we consider the origin of the fringes. Second, the light must in general be *monochromatic*, or nearly so. Especially is this true if the distances of  $M_1$  and  $M_2$  from  $G_1$  are appreciably different.

An extended source suitable for use with a Michelson interferometer may be obtained in any one of several ways. A sodium flame or a mercury arc, if large enough, may be used without the screen  $L$  shown in Fig. 13N. If the source is small, a ground-glass screen or a lens at  $L$  will extend the field of view. Looking at the mirror  $M_1$  through the plate  $G_1$ , one then sees the whole mirror filled with light. In order to obtain the fringes, the next step is to measure the distances of  $M_1$  and  $M_2$  to the back surface of  $G_1$  roughly with a millimeter scale and to move  $M_1$  until they are the same to within a few millimeters. The mirror  $M_2$  is now adjusted to be perpendicular to  $M_1$  by observing the images of a common pin, or any sharp point, placed between the source and  $G_1$ . Two pairs of images will be seen, one coming from reflection at the front surface of  $G_1$  and the other from reflection at its back surface. When the tilting screws on  $M_2$  are turned until one pair of images falls exactly on the other, the interference fringes should appear. When they first appear, the fringes will not be clear unless the eye is focused on or near the back mirror  $M_1$ , so the observer should look constantly at this mirror while searching for the fringes.

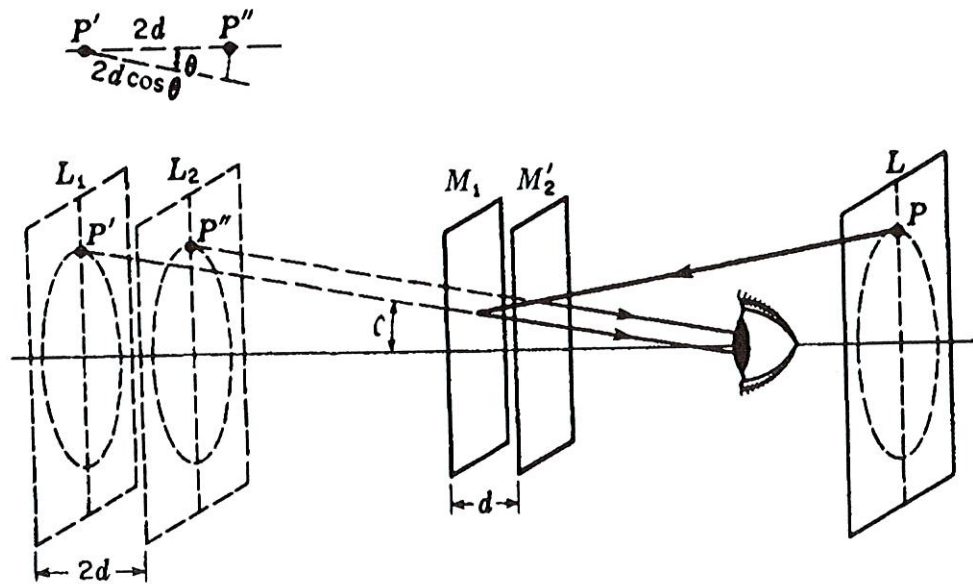


FIGURE 130  
Formation of circular fringes in the Michelson interferometer.

When they have been found, the adjusting screws should be turned in such a way as to continually increase the width of the fringes, and finally a set of concentric circular fringes will be obtained.  $M_2$  is then exactly perpendicular to  $M_1$  if the latter is at an angle of  $45^\circ$  with  $G_1$ .

### 13.9 CIRCULAR FRINGES

These are produced with monochromatic light when the mirrors are in exact adjustment and are the ones used in most kinds of measurement with the interferometer. Their origin can be understood by reference to the diagram of Fig. 130. Here the real mirror  $M_2$  has been replaced by its virtual image  $M_2'$  formed by reflection in  $G_1$ .  $M_2'$  is then parallel to  $M_1$ . Owing to the several reflections in the real interferometer, we may now think of the extended source as being at  $L$ , behind the observer, and as forming two virtual images  $L_1$  and  $L_2$  in  $M_1$  and  $M_2'$ . These virtual sources are coherent in that the phases of corresponding points in the two are exactly the same at all instants. If  $d$  is the separation  $M_1M_2'$ , the virtual sources will be separated by  $2d$ . When  $d$  is exactly an integral number of half wavelengths, i.e., the path difference  $2d$  equal to an integral number of whole wavelengths, all rays of light reflected normal to the mirrors will be in phase. Rays of light reflected at an angle, however, will in general not be in phase. The path difference between the two rays coming to the eye from corresponding points  $P'$  and  $P''$  is  $2d \cos \theta$ , as shown in the figure. The angle  $\theta$  is necessarily the same for the two rays when  $M_1$  is parallel to  $M_2'$  so that the rays are parallel. Hence when the eye is focused to receive parallel rays (a small telescope is more satisfactory here, especially for large values of  $d$ ) the rays will reinforce each other to produce maxima for those angles  $\theta$  satisfying the relation

$$2d \cos \theta = m\lambda \quad (13g)$$

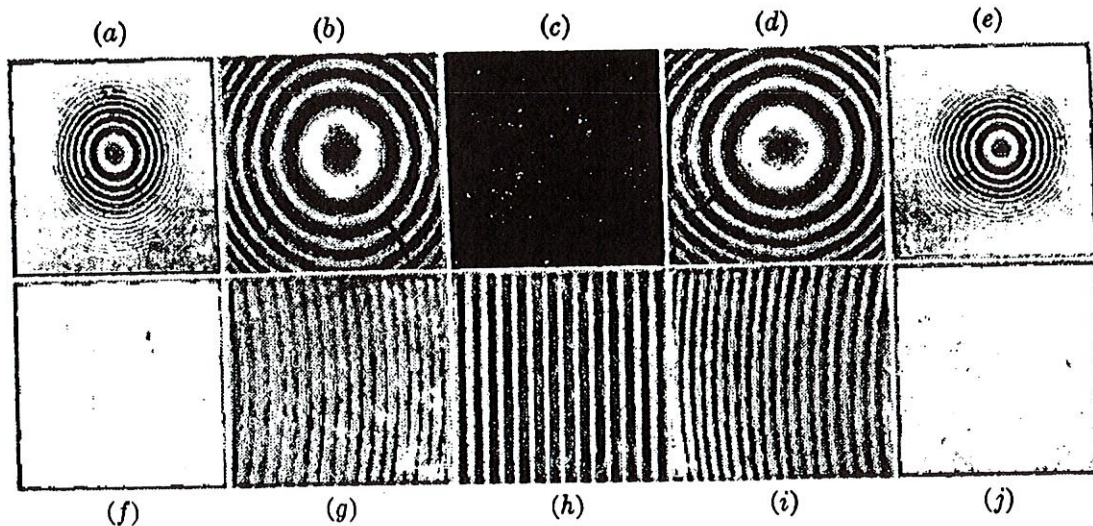


FIGURE 13P

Appearance of the various types of fringes observed in the Michelson interferometer. *Upper row, circular fringes. Lower row, localized fringes.* Path difference increases outward, in both directions, from the center.

Since for a given  $m$ ,  $\lambda$ , and  $d$  the angle  $\theta$  is constant, the maxima will lie in the form of circles about the foot of the perpendicular from the eye to the mirrors. By expanding the cosine, it can be shown from Eq. (13g) that the radii of the rings are proportional to the square roots of integers, as in the case of Newton's rings (Sec. 14.5). The intensity distribution across the rings follows Eq. (13b), in which the phase difference is given by

$$\delta = \frac{2\pi}{\lambda} 2d \cos \theta$$

Fringes of this kind, where parallel beams are brought to interference with a phase difference determined by the angle of inclination  $\theta$ , are often referred to as *fringes of equal inclination*. In contrast to the type to be described in the next section, this type may remain visible over very large path differences. The eventual limitation on the path difference will be discussed in Sec. 13.12.

The upper part of Fig. 13P shows how the circular fringes look under different conditions. Starting with  $M_1$  a few centimeters beyond  $M_2'$ , the fringe system will have the general appearance shown in (a) with the rings very closely spaced. If  $M_1$  is now moved slowly toward  $M_2'$  so that  $d$  is decreased, Eq. (13g) shows that a given ring, characterized by a given value of the order  $m$ , must decrease its radius because the product  $2d \cos \theta$  must remain constant. The rings therefore shrink and vanish at the center, a ring disappearing each time  $2d$  decreases by  $\lambda$ , or  $d$  by  $\lambda/2$ . This follows from the fact that at the center  $\cos \theta = 1$ , so that Eq. (13g) becomes

$$2d = m\lambda \quad (13h)$$

To change  $m$  by unity,  $d$  must change by  $\lambda/2$ . Now as  $M_1$  approaches  $M_2'$  the rings become more widely spaced, as indicated in Fig. 13P(b), until finally we reach a critical position where the central fringe has spread out to cover the whole field

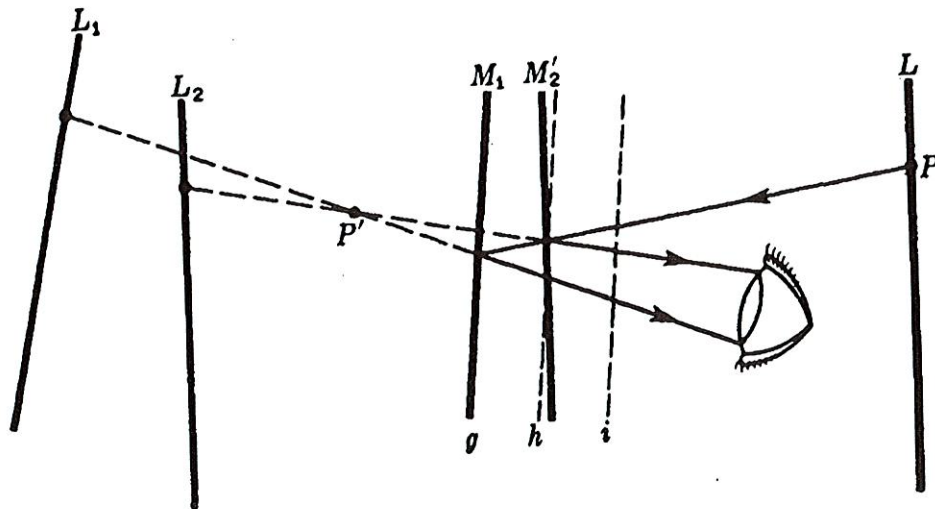


FIGURE 13Q  
The formation of fringes with inclined mirrors in the Michelson interferometer.

of view, as shown in (c). This happens when  $M_1$  and  $M_2'$  are exactly coincident, for it is clear that under these conditions the path difference is zero for all angles of incidence. If the mirror is moved still farther, it effectively passes through  $M_2'$ , and new widely spaced fringes appear, growing out from the center. These will gradually become more closely spaced as the path difference increases, as indicated in (d) and (e) of the figure.

### 13.10 LOCALIZED FRINGES

If the mirrors  $M_2'$  and  $M_1$  are not exactly parallel, fringes will still be seen with monochromatic light for path differences not exceeding a few millimeters. In this case the space between the mirrors is wedge-shaped, as indicated in Fig. 13Q. The two rays\* reaching the eye from a point  $P$  on the source are now no longer parallel, but appear to diverge from a point  $P'$  near the mirrors. For various positions of  $P$  on the extended source, it can be shown† that the path difference between the two rays remains constant but that the distance of  $P'$  from the mirrors changes. If the angle between the mirrors is not too small, however, the latter distance is never great, and hence, in order to see these fringes clearly, the eye must be focused on or near the rear mirror  $M_1$ . The localized fringes are practically straight because the variation of the path difference across the field of view is now due primarily to the variation of the thickness of the "air film" between the mirrors. With a wedge-shaped film, the locus of points of equal thickness is a straight line parallel to the edge of the wedge. The

\* When the term "ray" is used, here and elsewhere in discussing interference phenomena, it merely indicates the direction of the *perpendicular to a wave front* and is in no way to suggest an infinitesimally narrow pencil of light.

† R. W. Ditchburn, "Light," 2d ed., paperback, John Wiley and Sons, Inc., New York, 1963.

fringes are not exactly straight, however, if  $d$  has an appreciable value, because there is also some variation of the path difference with angle. They are in general curved and are always convex toward the thin edge of the wedge. Thus, with a certain value of  $d$ , we might observe fringes shaped like those of Fig. 13P(*g*).  $M_1$  could then be in a position such as *g* of Fig. 13Q. If the separation of the mirrors is decreased, the fringes will move to the left across the field, a new fringe crossing the center each time  $d$  changes by  $\lambda/2$ . As we approach zero path difference, the fringes become straighter, until the point is reached where  $M_1$  actually intersects  $M_2'$ , when they are perfectly straight, as in (*h*). Beyond this point, they begin to curve in the opposite direction, as shown in (*i*). The blank fields (*f*) and (*j*) indicate that this type of fringe cannot be observed for large path differences. Because the principal variation of path difference results from a change of the thickness  $d$ , these fringes have been termed *fringes of equal thickness*.

### 13.11 WHITE-LIGHT FRINGES

If a source of white light is used, no fringes will be seen at all except for a path difference so small that it does not exceed a few wavelengths. In observing these fringes, the mirrors are tilted slightly as for localized fringes, and the position of  $M_1$  is found where it intersects  $M_2'$ . With white light there will then be observed a central dark fringe, bordered on either side by 8 or 10 colored fringes. This position is often rather troublesome to find using white light only. It is best located approximately beforehand by finding the place where the localized fringes in monochromatic light become straight. Then a very *slow* motion of  $M_1$  through this region, using white light, will bring these fringes into view.

The fact that only a few fringes are observed with white light is easily accounted for when we remember that such light contains all wavelengths between 400 and 750 nm. The fringes for a given color are more widely spaced the greater the wavelength. Thus the fringes in different colors will only coincide for  $d = 0$ , as indicated in Fig. 13R. The solid curve represents the intensity distribution in the fringes for green light, and the broken curve that for red light. Clearly, only the central fringe will be uncolored, and the fringes of different colors will begin to separate at once on either side, producing various impure colors which are not the saturated spectral colors. After 8 or 10 fringes, so many colors are present at a given point that the resultant color is essentially white. Interference is still occurring in this region, however, because a spectroscope will show a continuous spectrum with dark bands at those wavelengths for which the condition for destructive interference is fulfilled. White-light fringes are also observed in all the other methods of producing interference described above, if white light is substituted for monochromatic light. They are particularly important in the Michelson interferometer, where they may be used to locate the position of zero path difference, as we shall see in Sec. 13.13.

An excellent reproduction in color of these white-light fringes is given in one of Michelson's books.\* The fringes in three different colors are also shown separately

\* A. A. Michelson, "Light Waves and Their Uses," plate II, University of Chicago Press, Chicago, 1906.

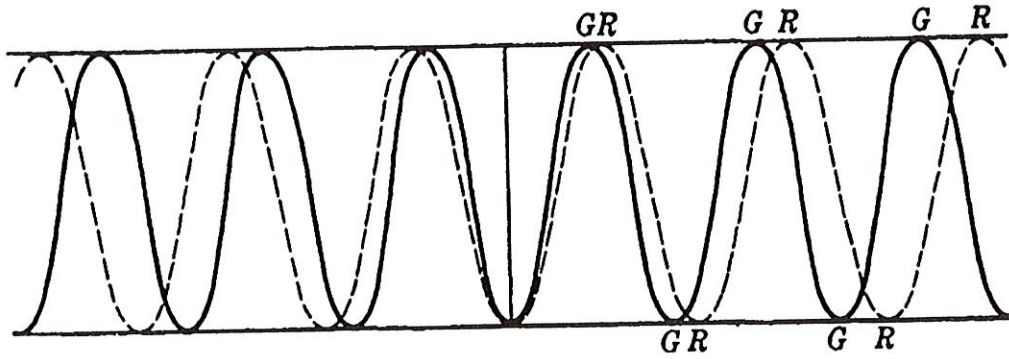


FIGURE 13R  
The formation of white-light fringes with a dark fringe at the center.

and a study of these in connection with the white-light fringes is instructive as showing the origin of the various impure colors in the latter.

It was stated above that the central fringe in the white-light system, i.e., that corresponding to zero path difference, is black when observed in the Michelson interferometer. One would ordinarily expect this fringe to be white, since the two beams should be in phase with each other for any wavelength at this point, and in fact this is the case in the fringes formed with the other arrangements, such as the biprism. In the present case, however, it will be seen by referring to Fig. 13N that while ray 1 undergoes an internal reflection in the plate  $G_1$ , ray 2 undergoes an external reflection, with a consequent change of phase [see Eq. (14d)]. Hence the central fringe is black if the back surface of  $G_1$  is unsilvered. If it is silvered, the conditions are different and the central fringe may be white.

### 13.12 VISIBILITY OF THE FRINGES

There are three principal types of measurement that can be made with the interferometer: (1) width and fine structure of spectrum lines, (2) lengths or displacements in terms of wavelengths of light, and (3) refractive indices. As explained in the preceding section, when a certain spread of wavelengths is present in the light source, the fringes become indistinct and eventually disappear as the path difference is increased. With white light they become invisible when  $d$  is only a few wavelengths, whereas the circular fringes obtained with the light of a single spectrum line can still be seen after the mirror has been moved several centimeters. Since no line is perfectly sharp, however, the different component wavelengths produce fringes of slightly different spacing, and hence there is a limit to the usable path difference even in this case. For the measurements of length to be described below, Michelson tested the lines from various sources and concluded that a certain red line in the spectrum of cadmium was the most satisfactory. He measured the *visibility*, defined as

$$V = \frac{I_{\max} - I_{\min}}{I_{\max} + I_{\min}} \quad (13i)$$